



planck

Planck observations of CMB polarisation



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- **questions :**

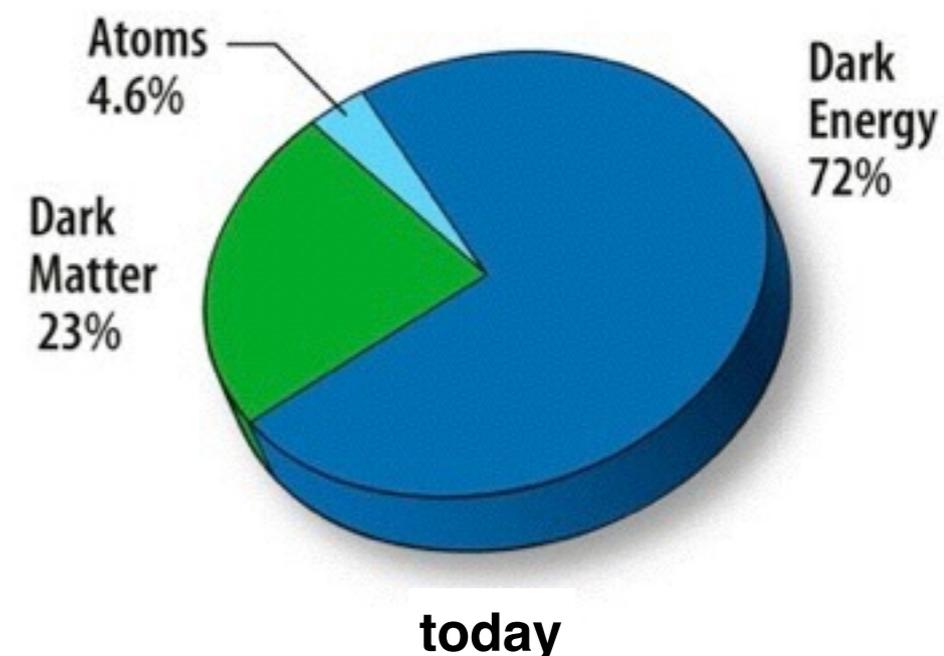
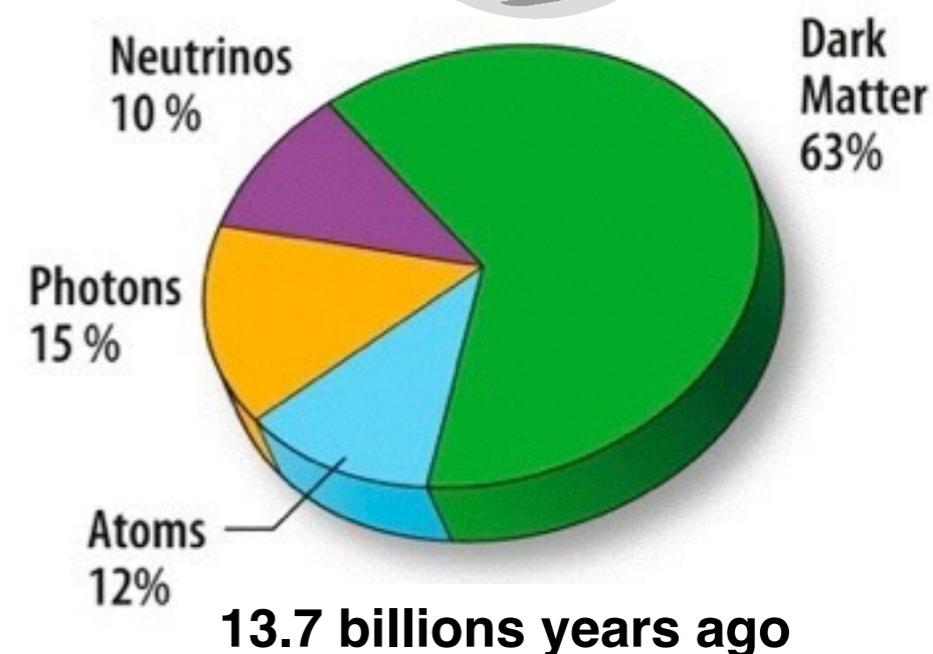
- How old is the Universe ?
- What's its composition ?
- How could we describe its evolution ?

- **that leads to more questions :**

- What is dark matter ?
- What is Dark Energy ?
- ...

- **informations :**

- study matter distribution at different time during Univers evolution



Big-Bang model



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10^{-32} seconds

1 second

100 seconds

380 000 years

300–500 million years

Billions of years

13.8 billion years

Beginning
of the
Universe



Inflation

Accelerated expansion
of the Universe

Formation of light and matter

Light and matter are coupled

Dark matter evolves
independently: it starts
clumping and forming
a web of structures

Light and matter separate

- Protons and electrons
form atoms
- Light starts travelling
freely: it will become the
Cosmic Microwave
Background (CMB)

Dark ages

Atoms start feeling
the gravity of the
cosmic web of dark
matter

First stars

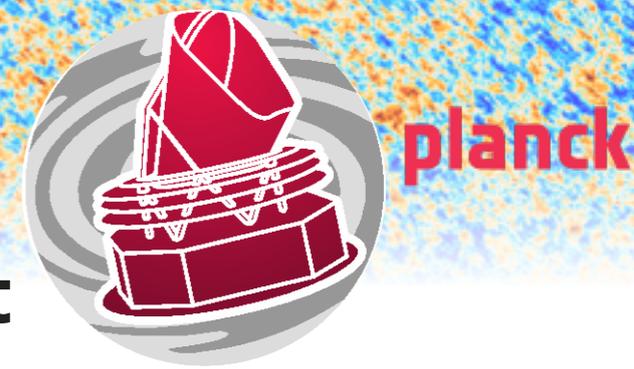
The first stars and
galaxies form in the
densest knots of the
cosmic web

Galaxy evolution

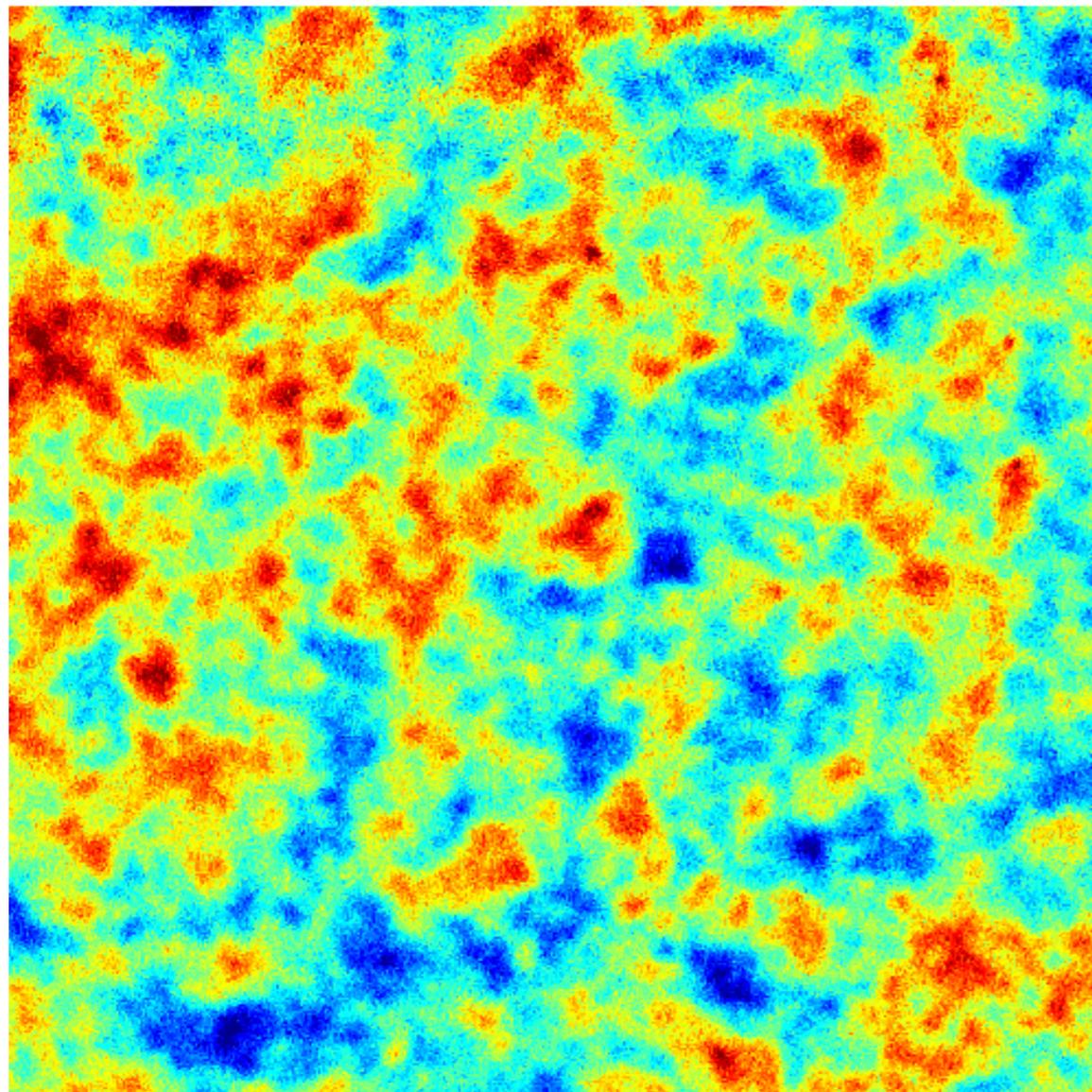
The present Universe

CMB anisotropies

match the matter distribution in the Universe at very early times



zoom $10 \times 10^\circ$



-300  300 μK
(30.0, 40.0) Galactic

- **Gaussian variable correlated in real space**
- **decorrelated in Fourier domain**

$$\begin{aligned} c(\theta) &= \langle T(\hat{n}) \times T(\hat{n} + \theta) \rangle \\ &= \sum_{\ell} C_{\ell} (2\ell + 1) P_{\ell}(\cos\theta) \end{aligned}$$

- **encodes information on the primordial fluctuations**

angular power spectrum



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toy-model

- **causal physics**

$$\theta_H \simeq 1.2^\circ$$

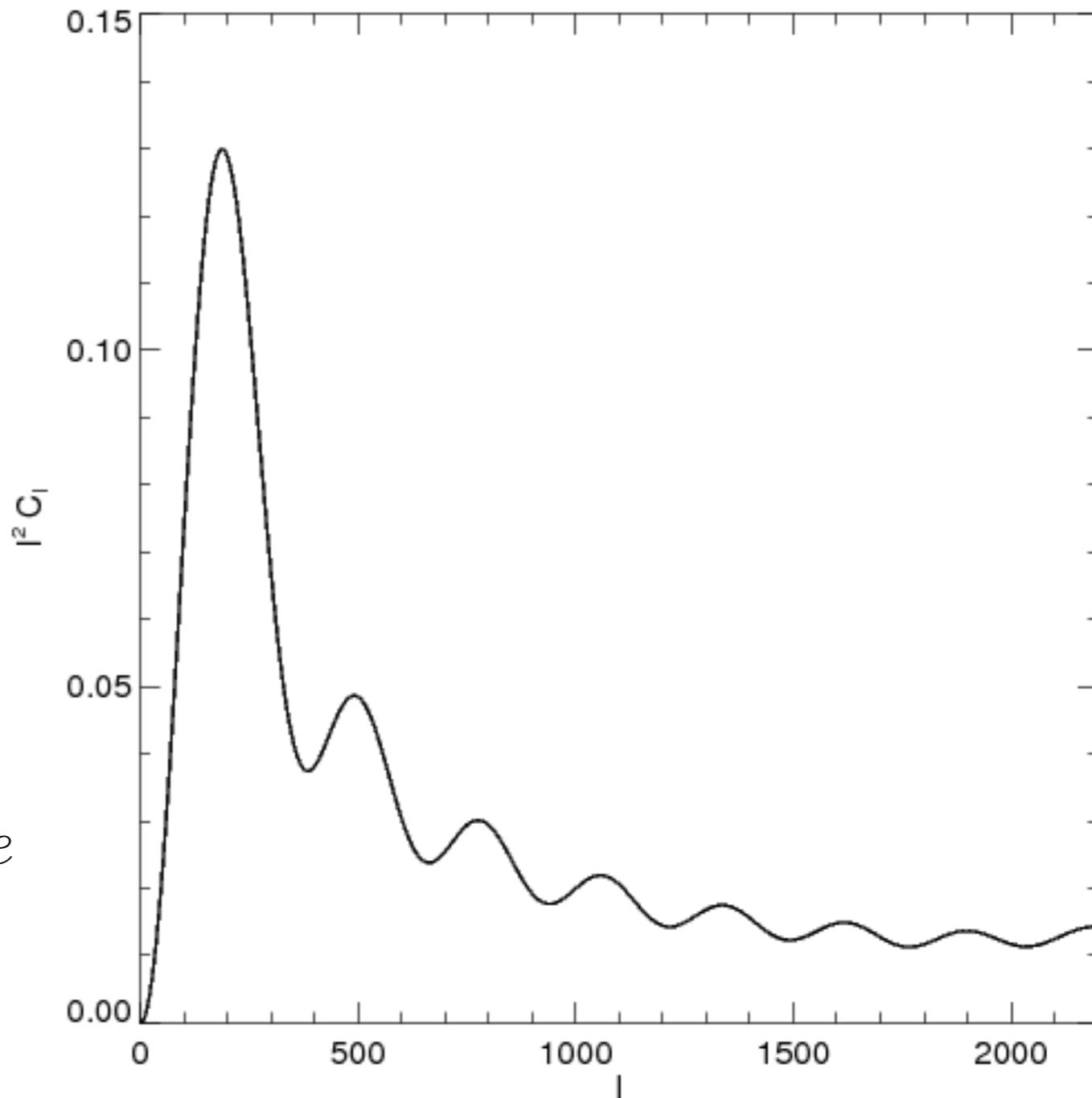
- **auto-correlation**

$$c(\theta) = \begin{cases} (\theta - \theta_H)^2 & \theta < \theta_H \\ 0 & \theta > \theta_H \end{cases}$$

- **on the sphere**

$$l(l+1)C_l \simeq \left(1 - \frac{\sin l\theta_H}{l\theta_H}\right) \otimes \text{sphere}$$

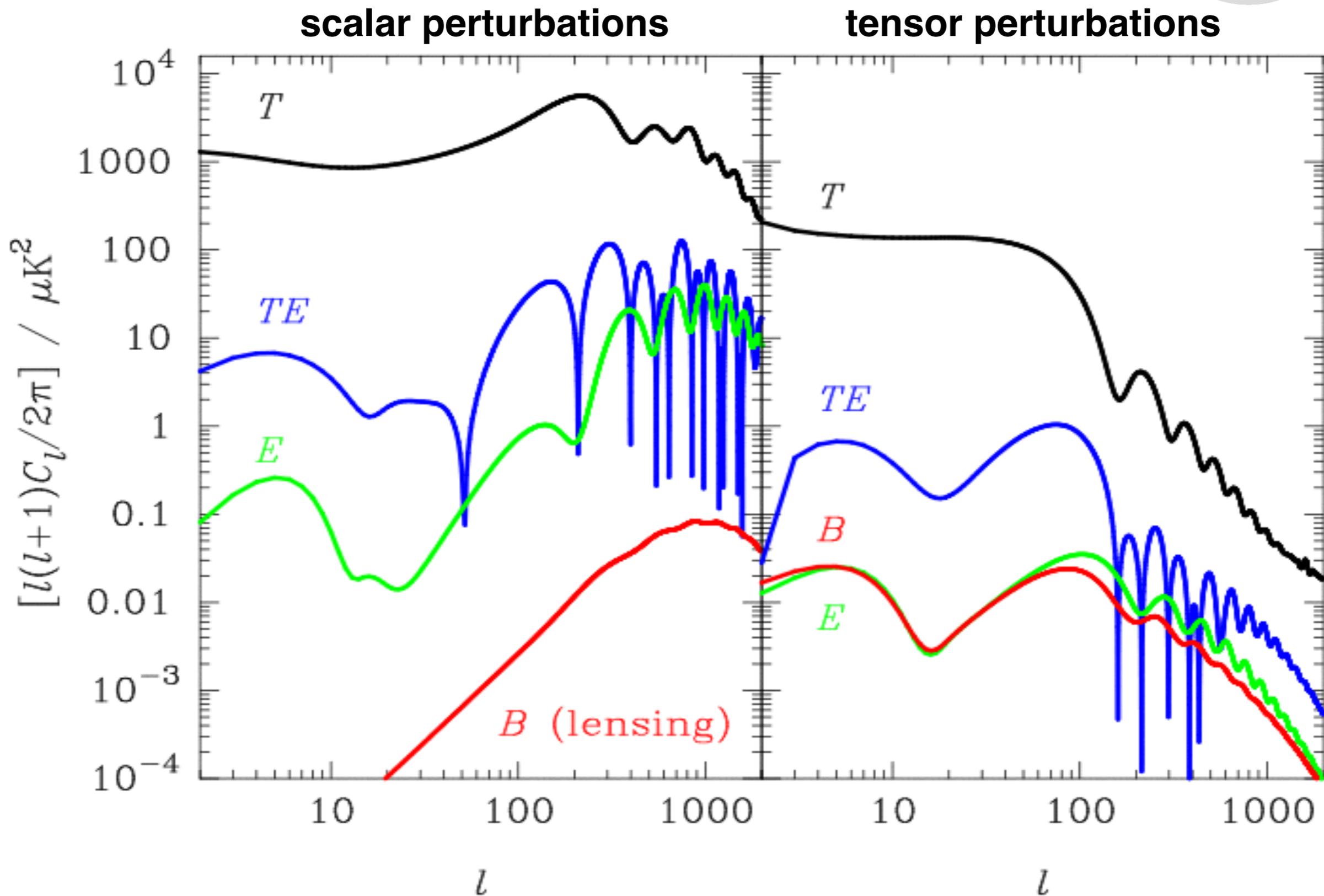
$$l = 220 \leftrightarrow \theta \simeq 1.2^\circ$$



Cosmological model



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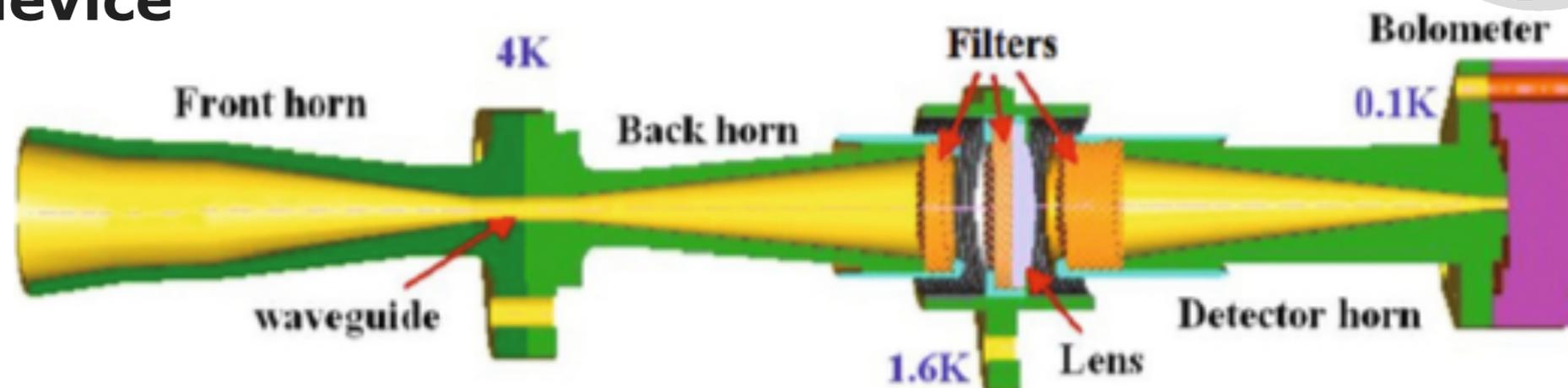


Planck polarised detectors

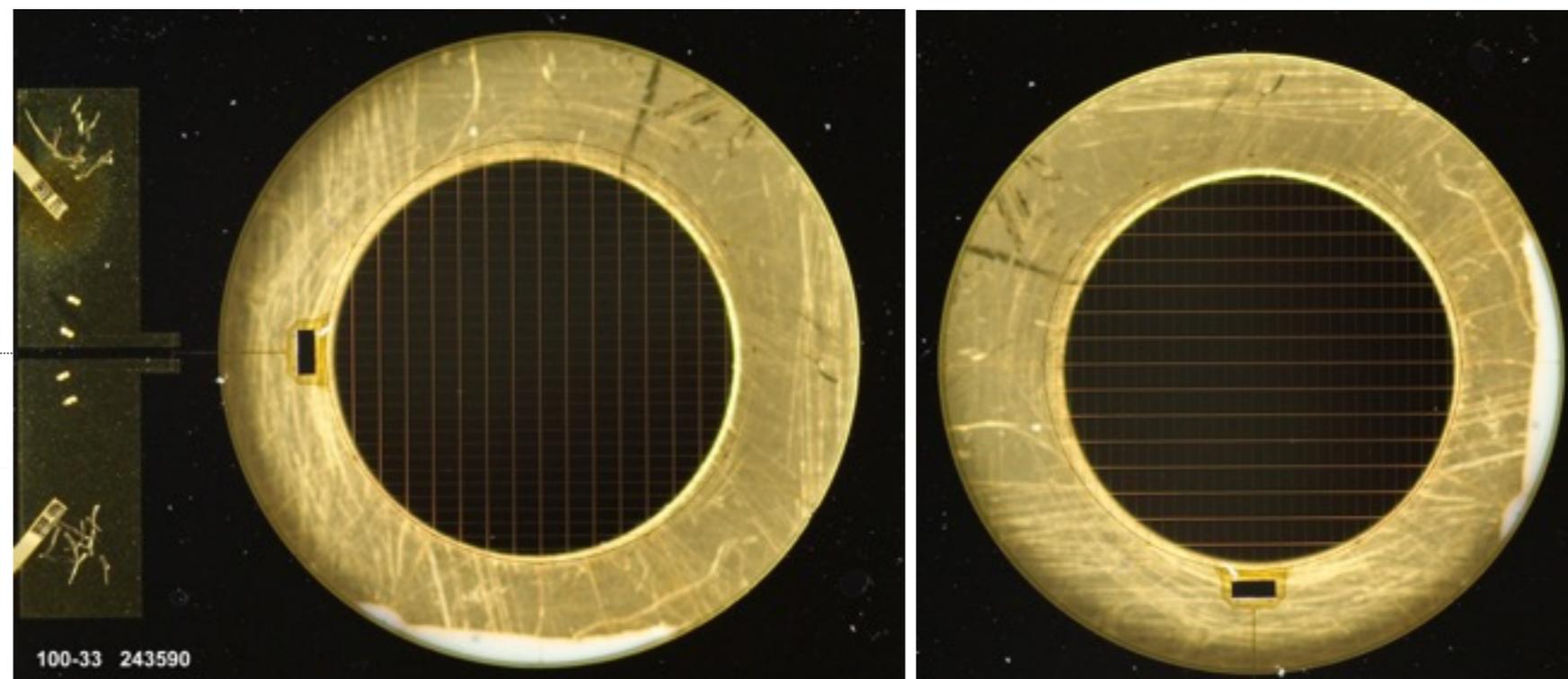
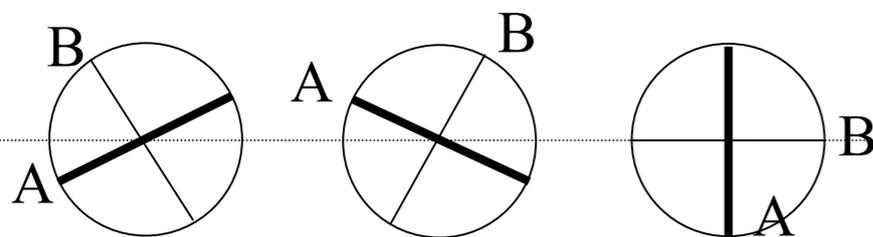


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- optical device

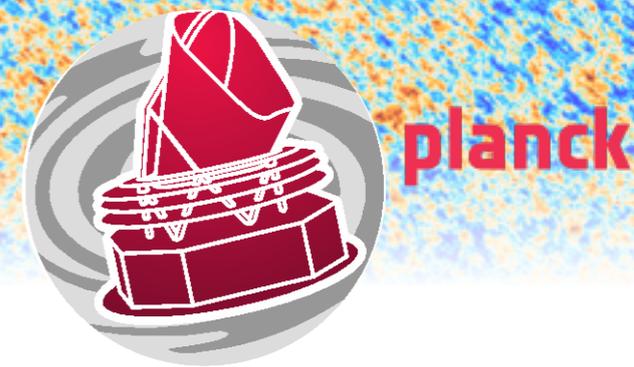


- Bolometers



$$m_t = I(\vec{n}) + \rho \left[Q(\vec{n}) \cdot \cos(2\psi) + U(\vec{n}) \cdot \sin(2\psi) \right]$$

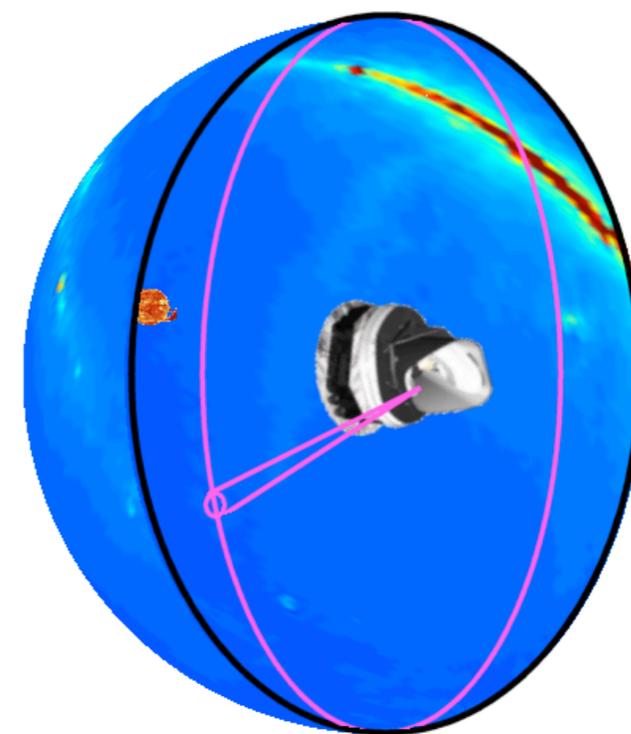
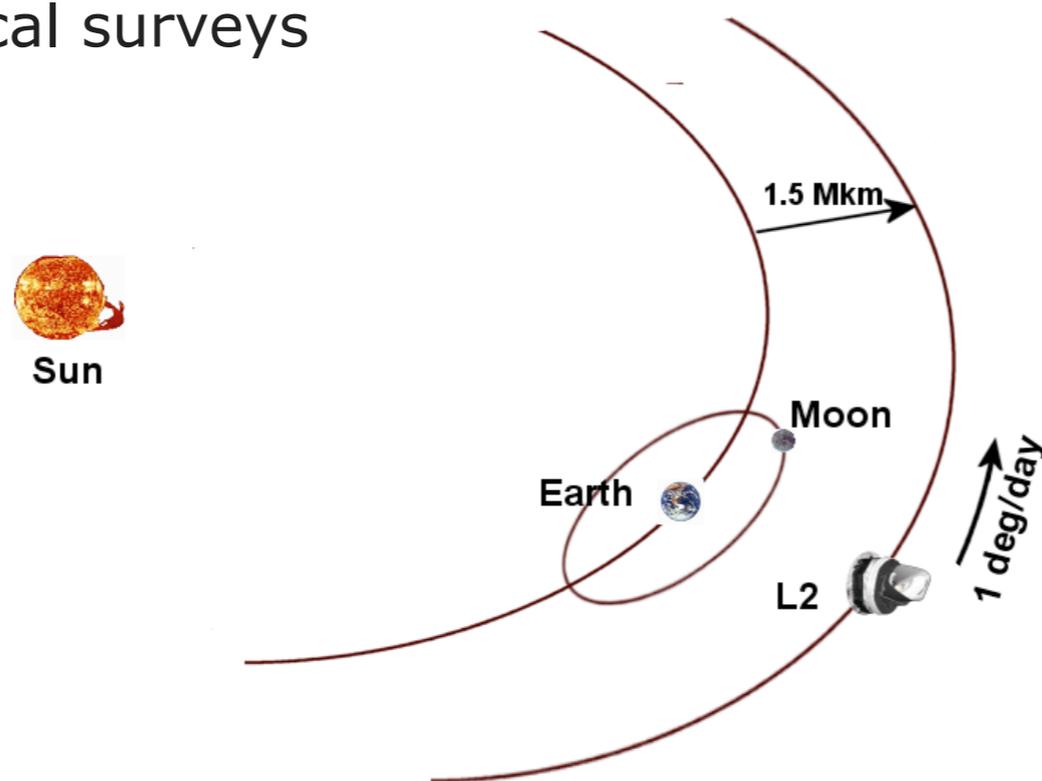
Polarisation measurement



- we need different angles to measure **I,Q,U**

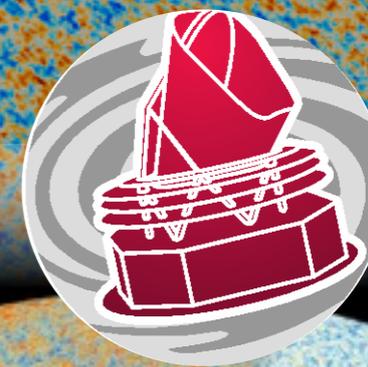
$$m_t = I(\vec{n}) + \rho \left[Q(\vec{n}) \cdot \cos(2\psi) + U(\vec{n}) \cdot \sin(2\psi) \right]$$

- **Planck scanning strategy is such that we have**
 - one orientation of the focal plane / sky pixel / survey
 - 2 pairs of identical surveys

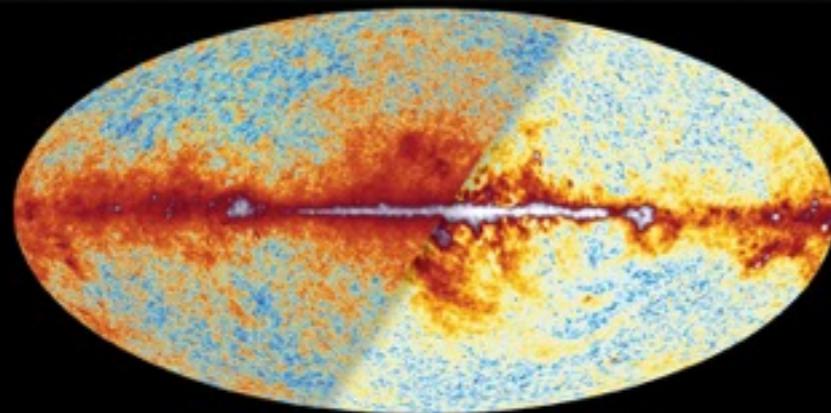


- **we combine detectors at 90deg**
 - need to have very precise inter-calibration to avoid I,Q,U mixing
 - Intensity signal is ~ 100 to 1000 times larger than polarisation

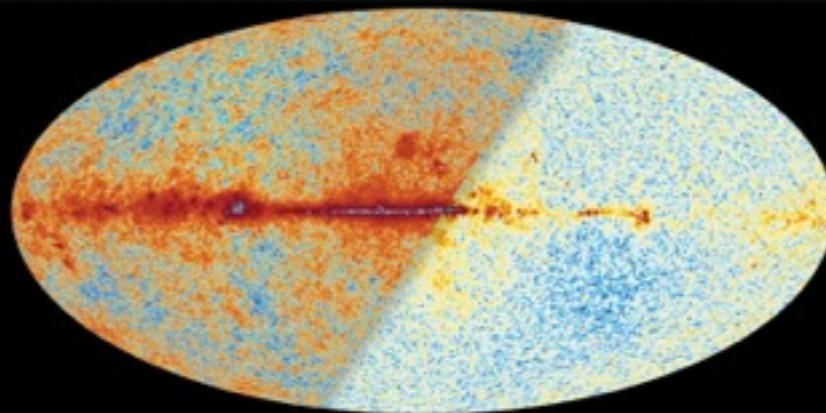
the sky as seen by Planck



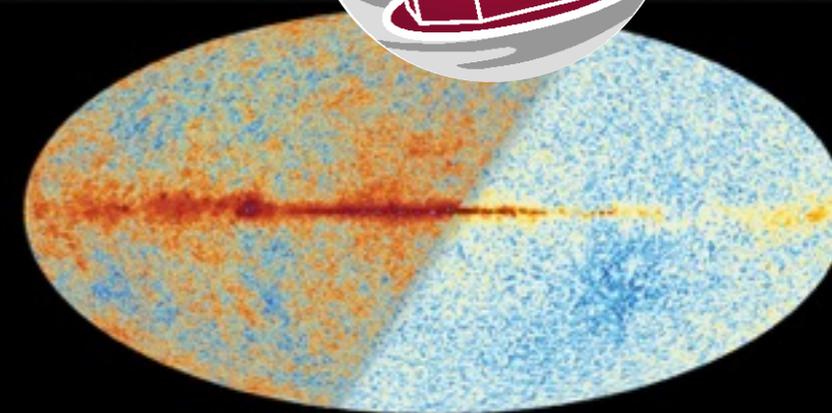
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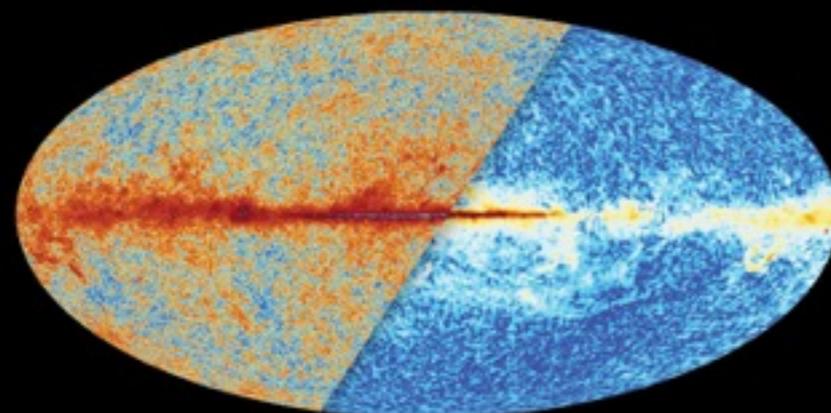
30 GHz



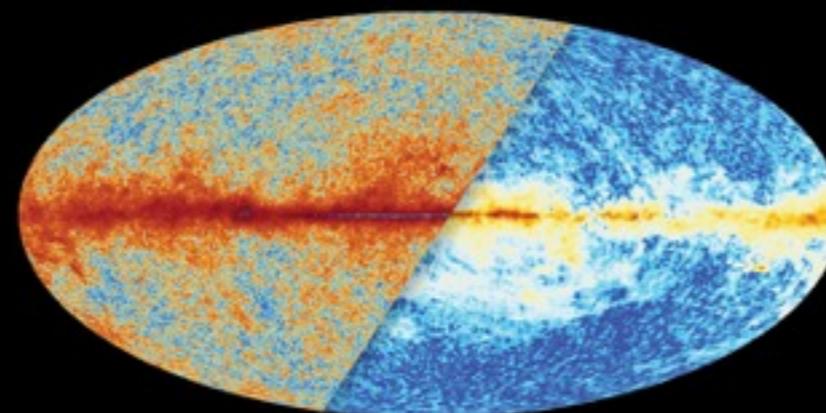
44 GHz



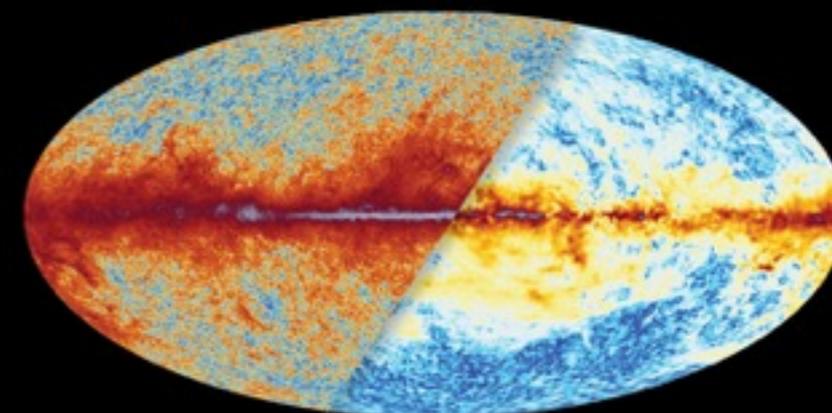
70 GHz



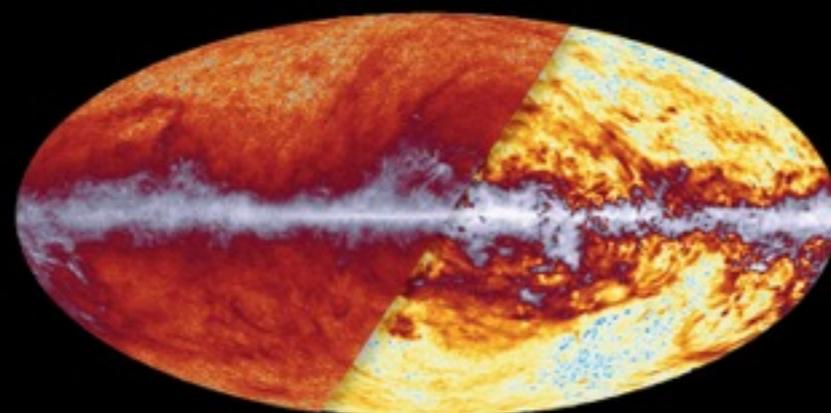
100 GHz



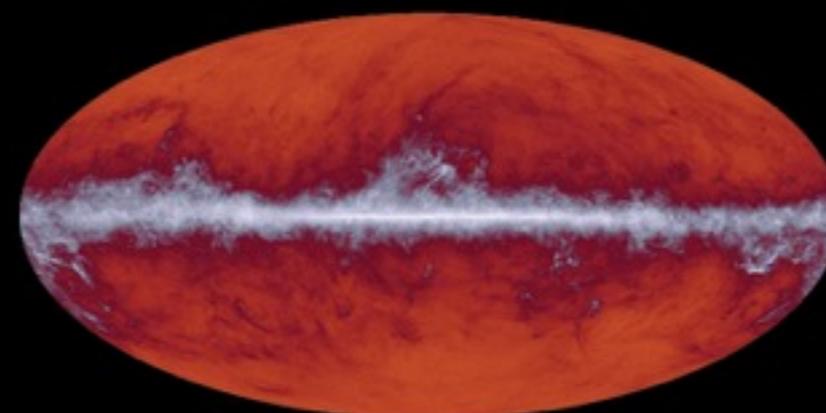
143 GHz



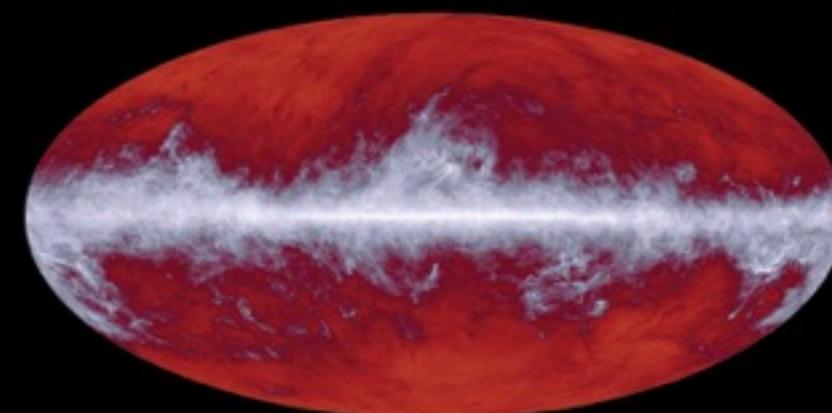
217 GHz



353 GHz



545 GHz

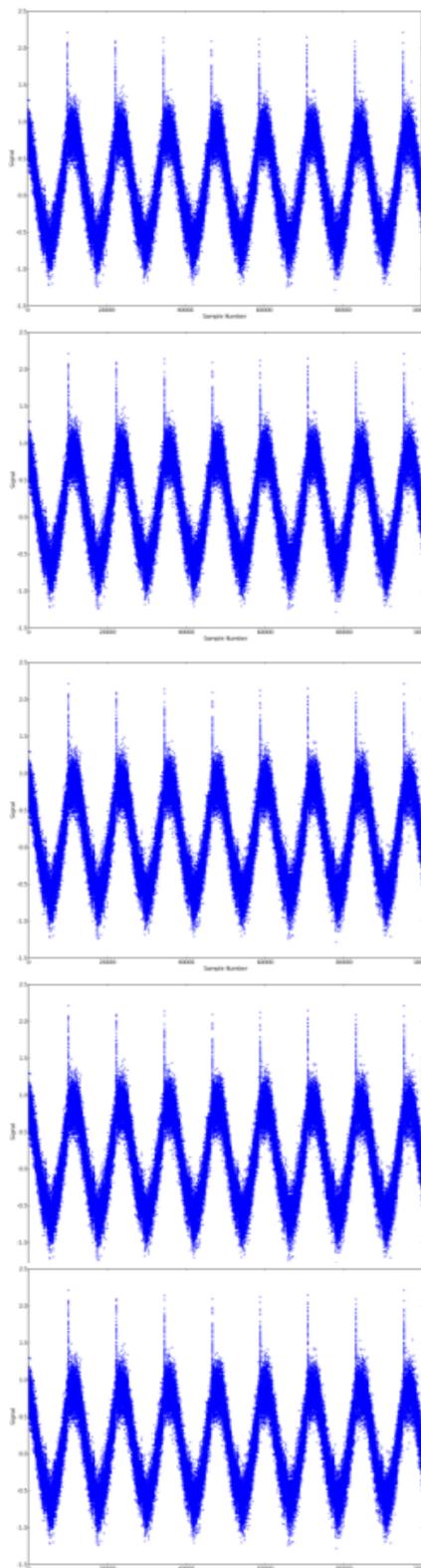


857 GHz

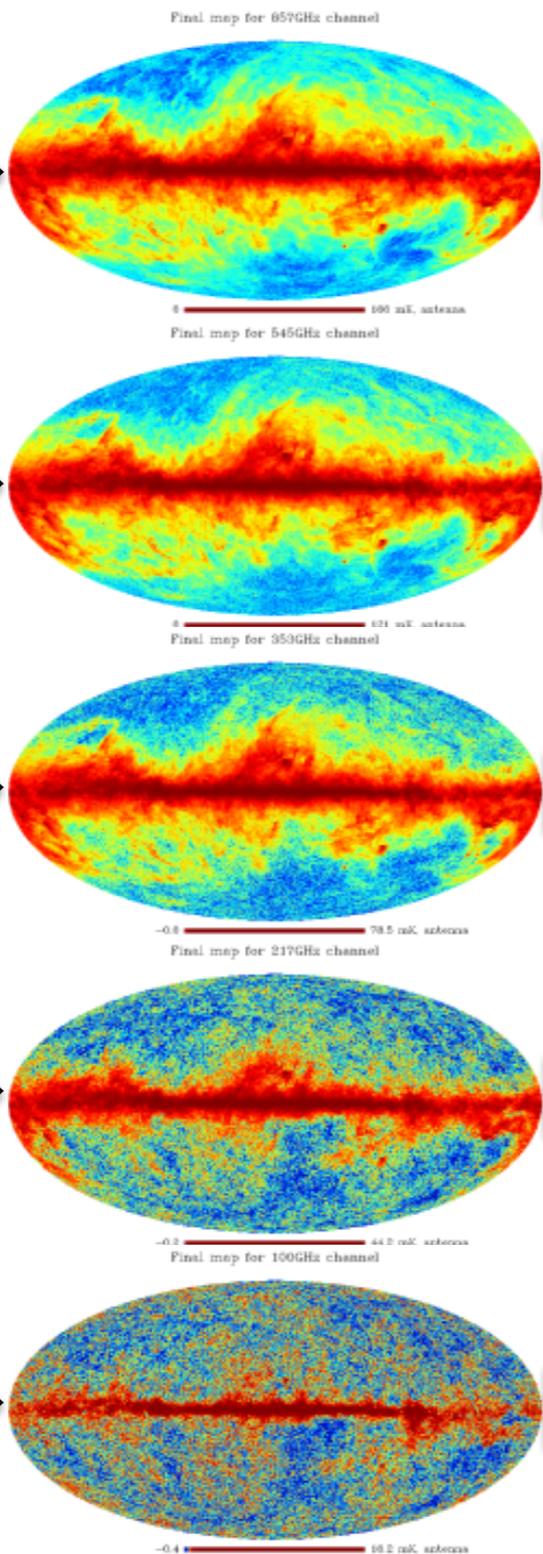
CMB data treatment



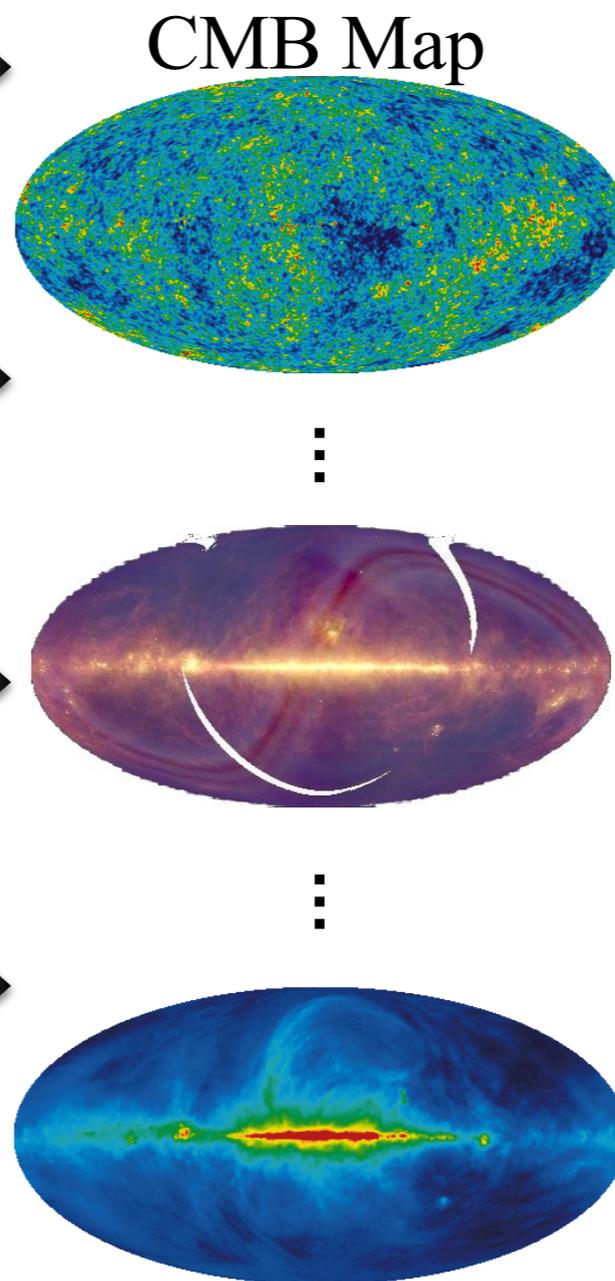
data



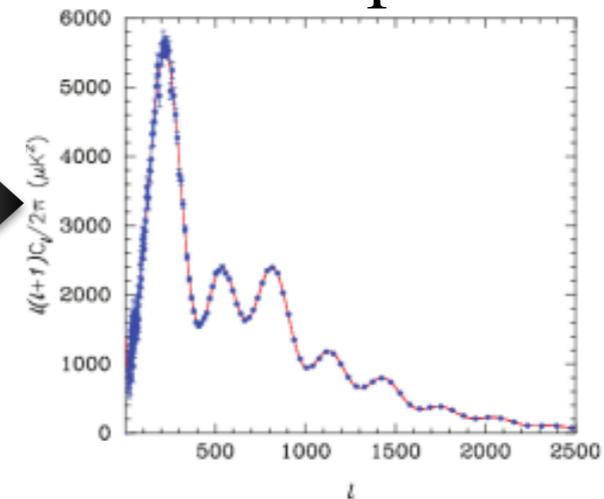
8 frequency maps



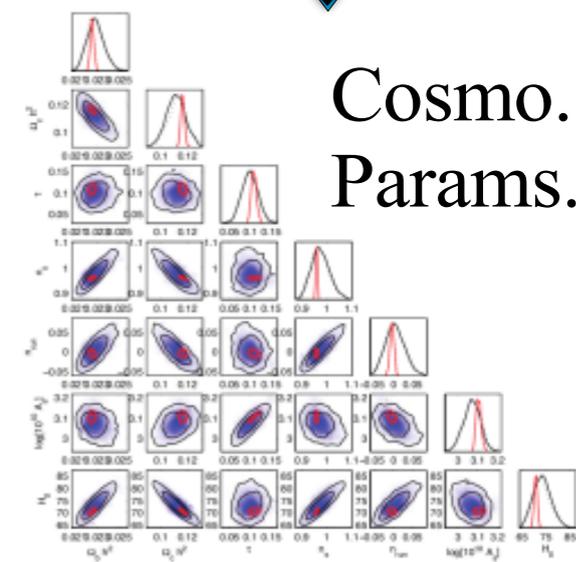
components maps



CMB Spectra



Cosmo.
Params.

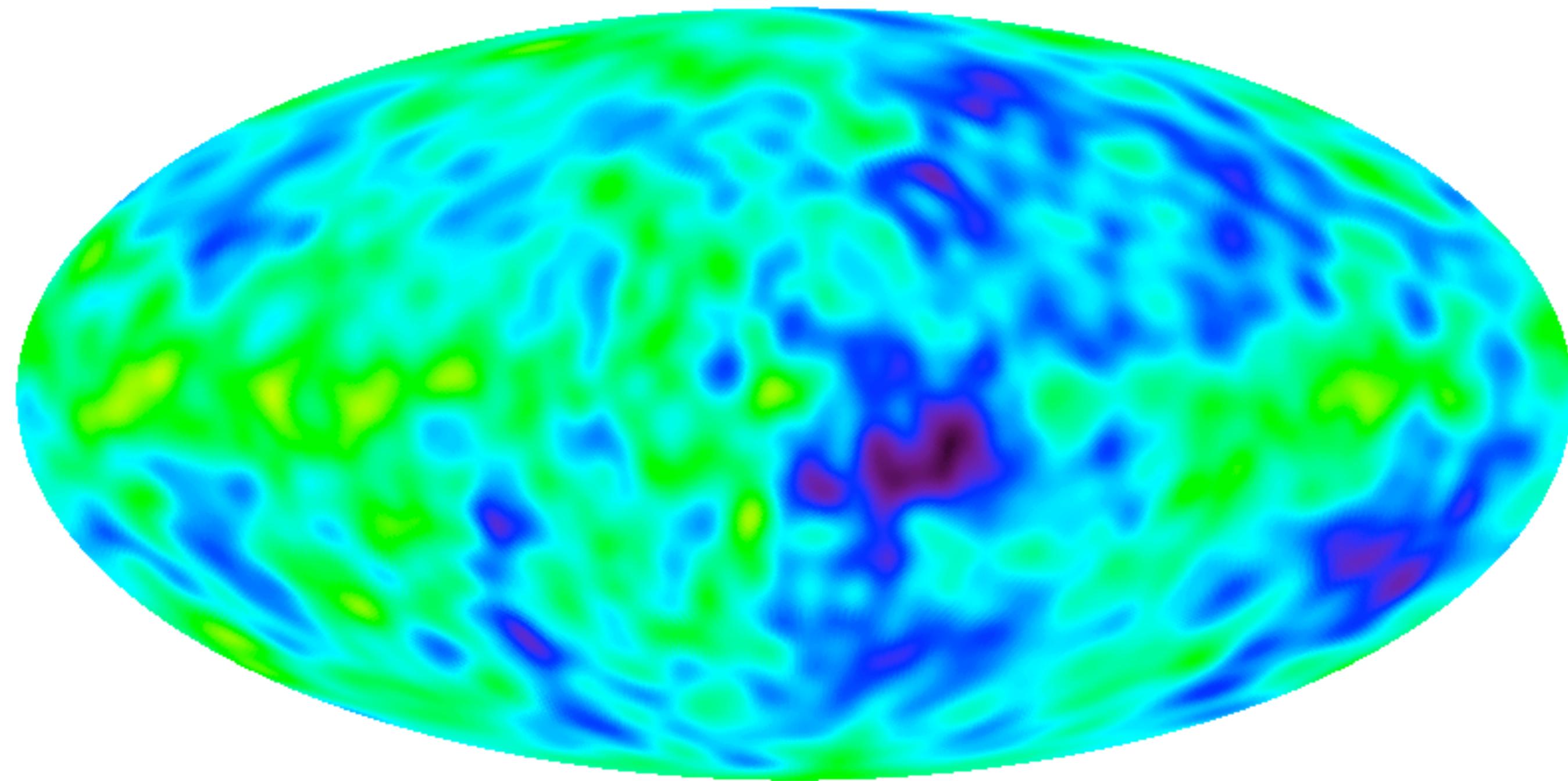


COBE

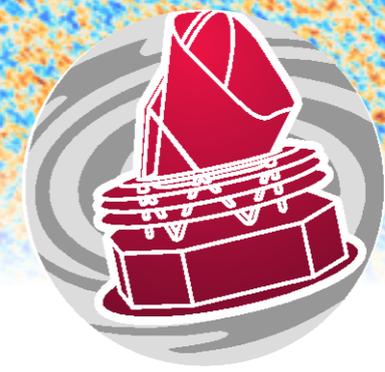


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1993

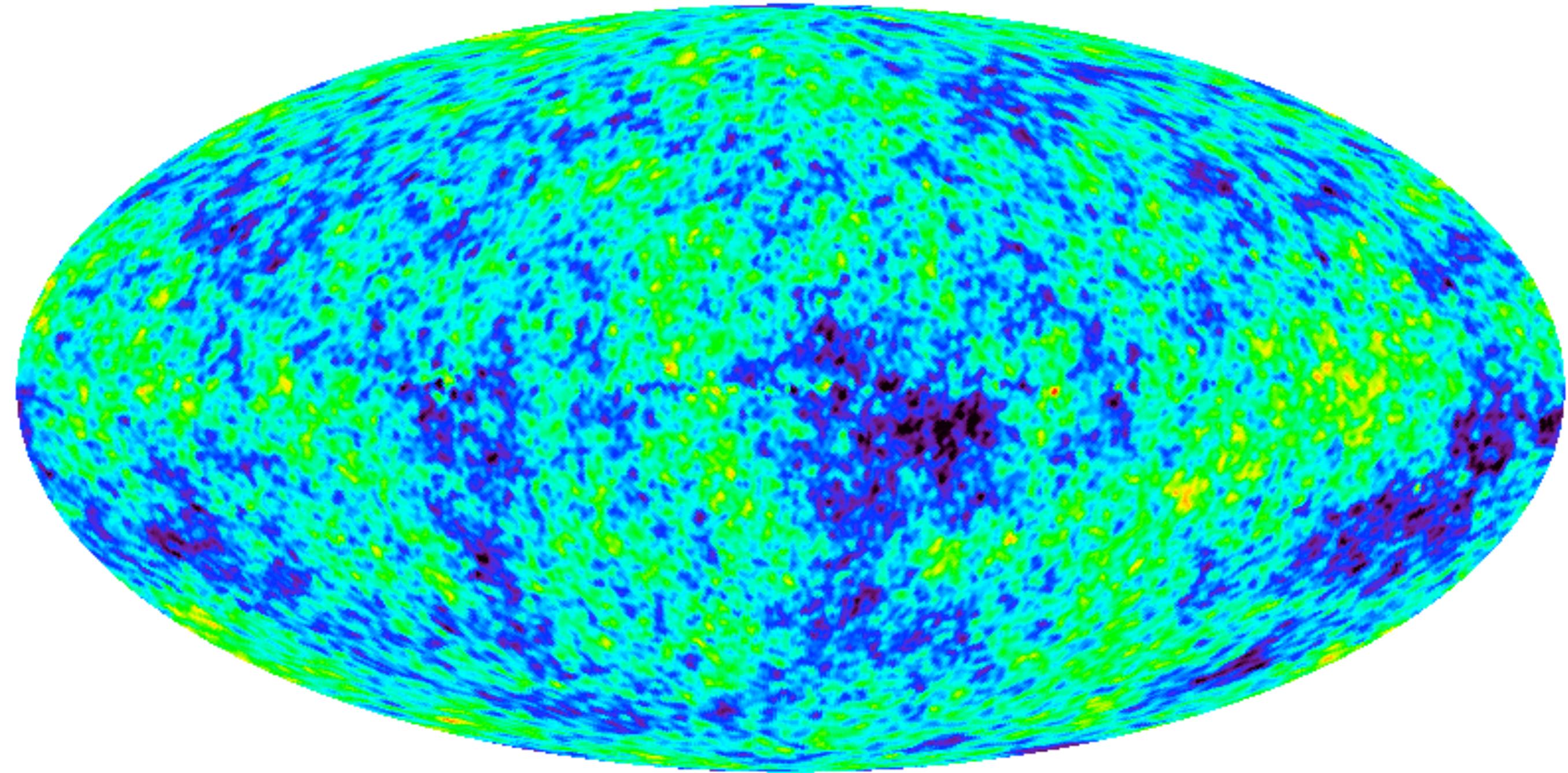


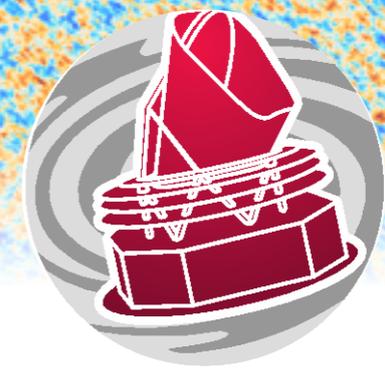
WMAP



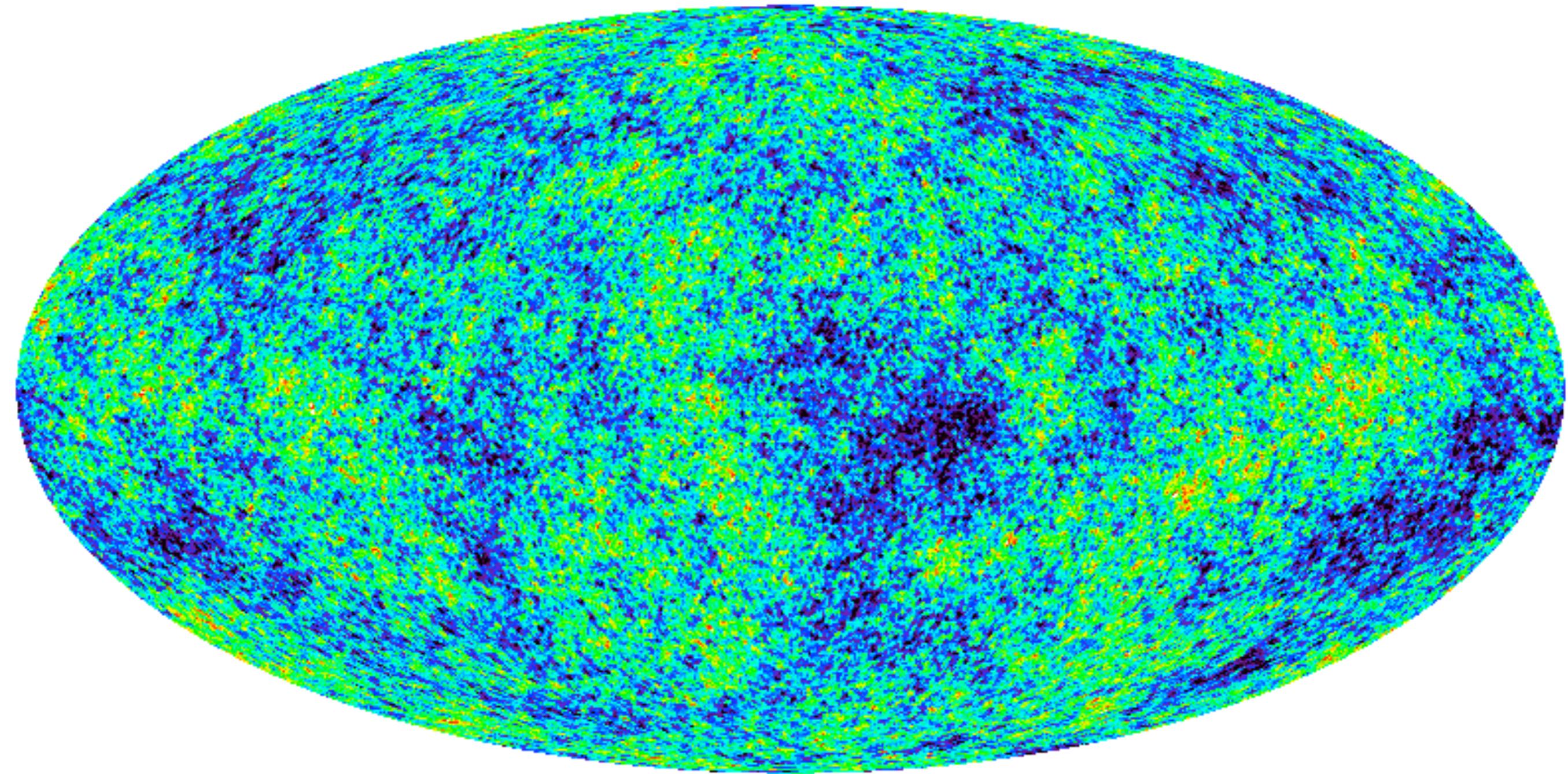
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2003





2013

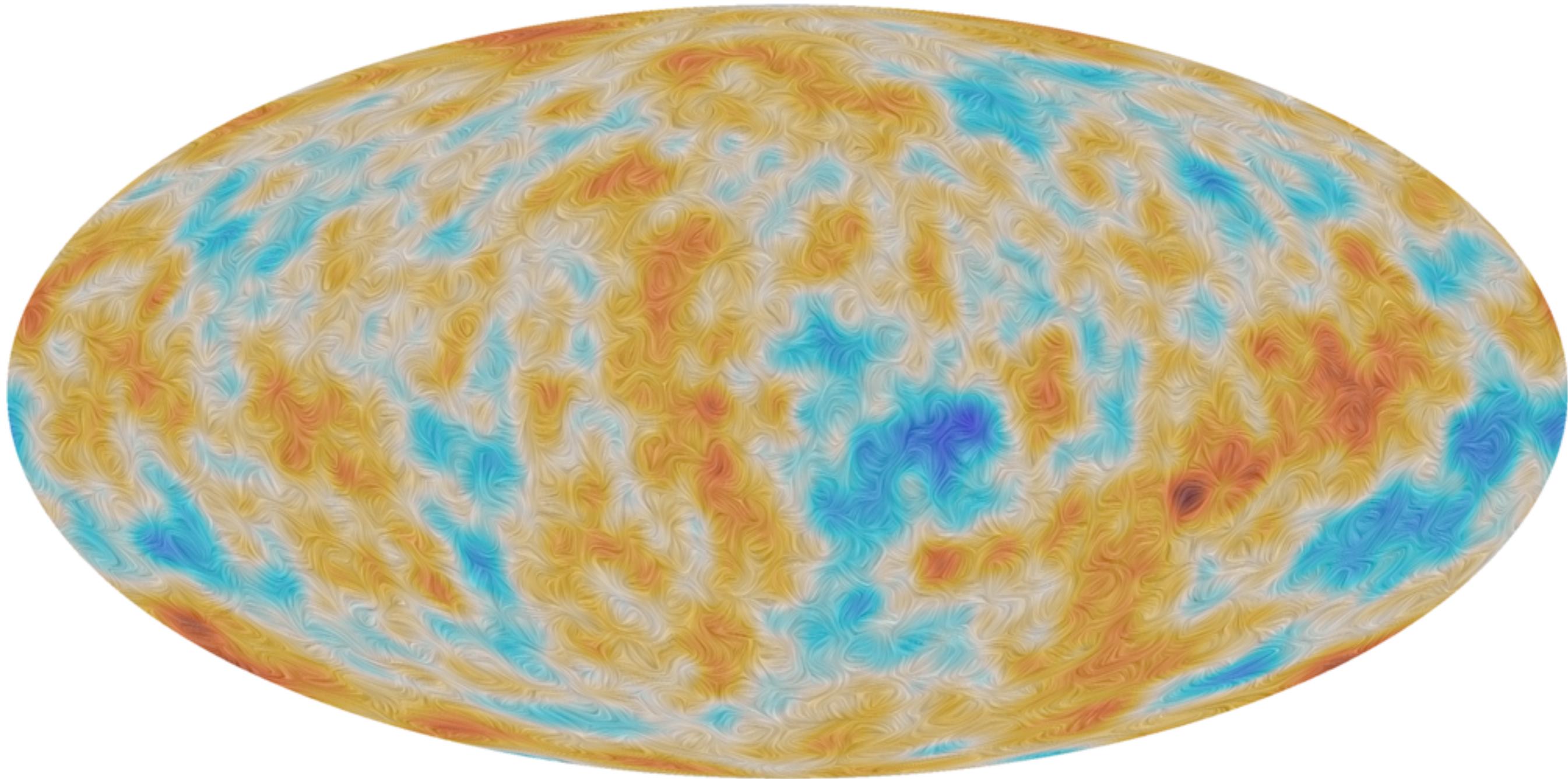


Planck



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2015

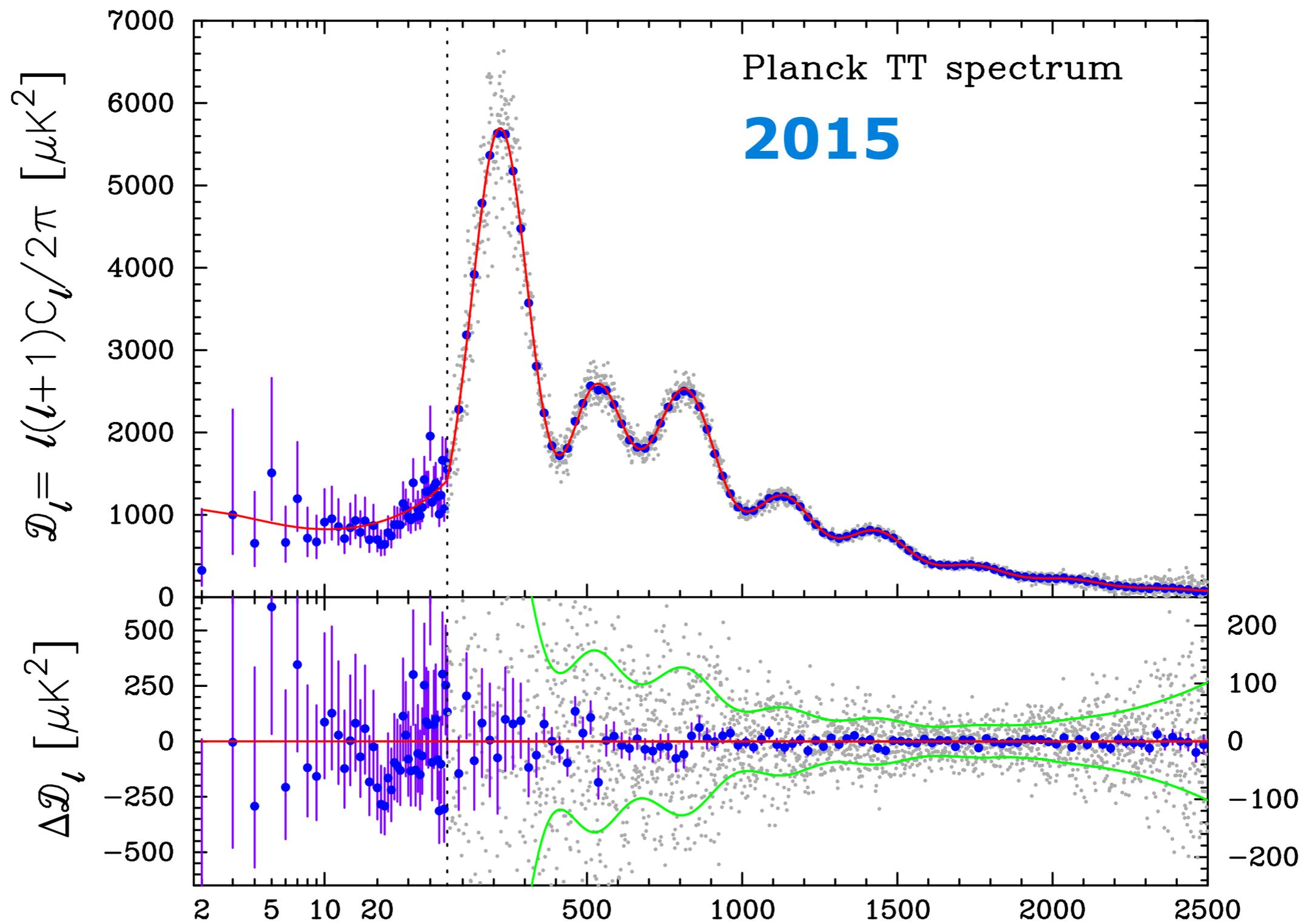


polarization

temperature power spectrum



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[Planck Collaboration XIII, *Planck* 2015 results]

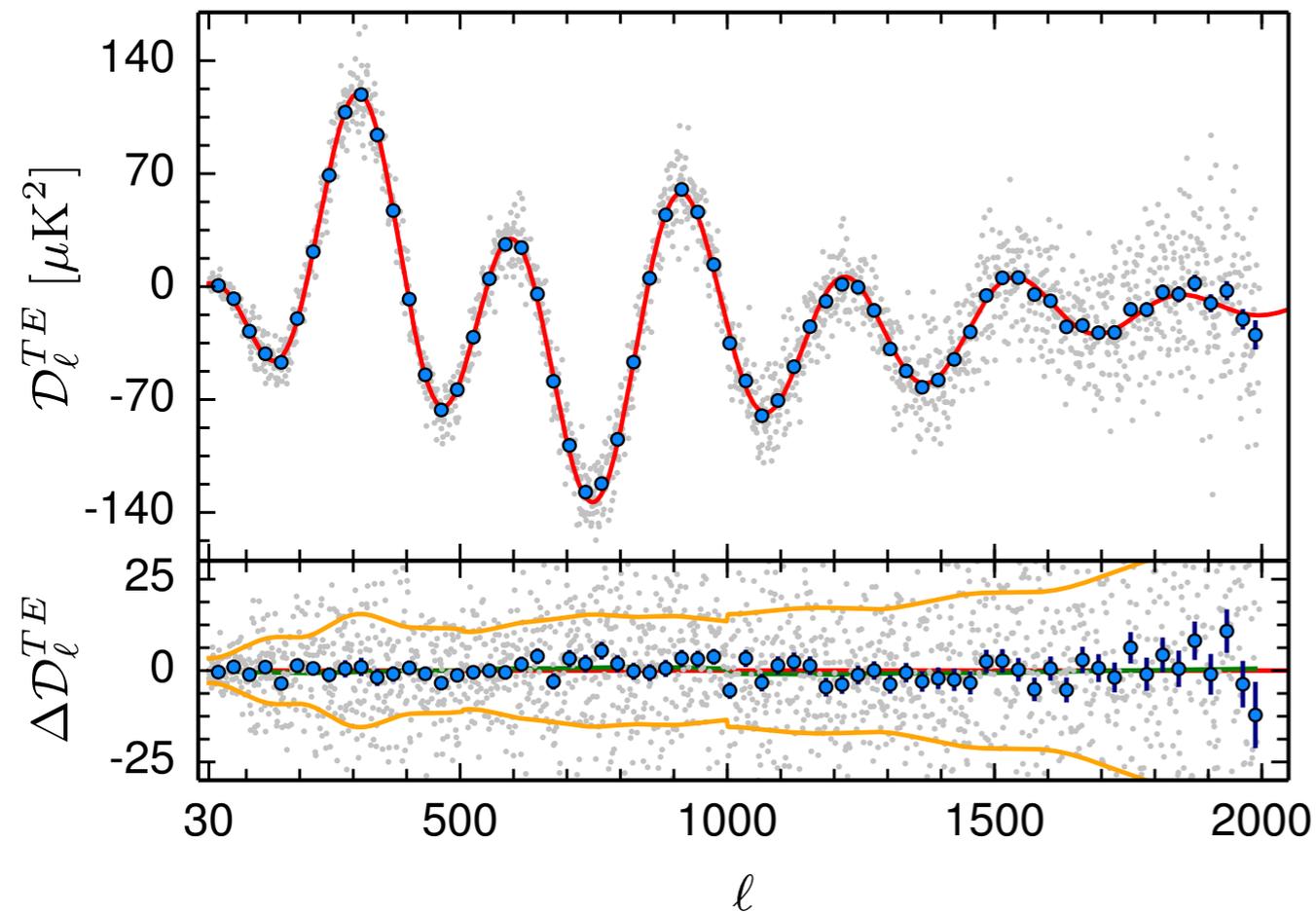
Planck observations of CMB polarisation

polarized power spectrum

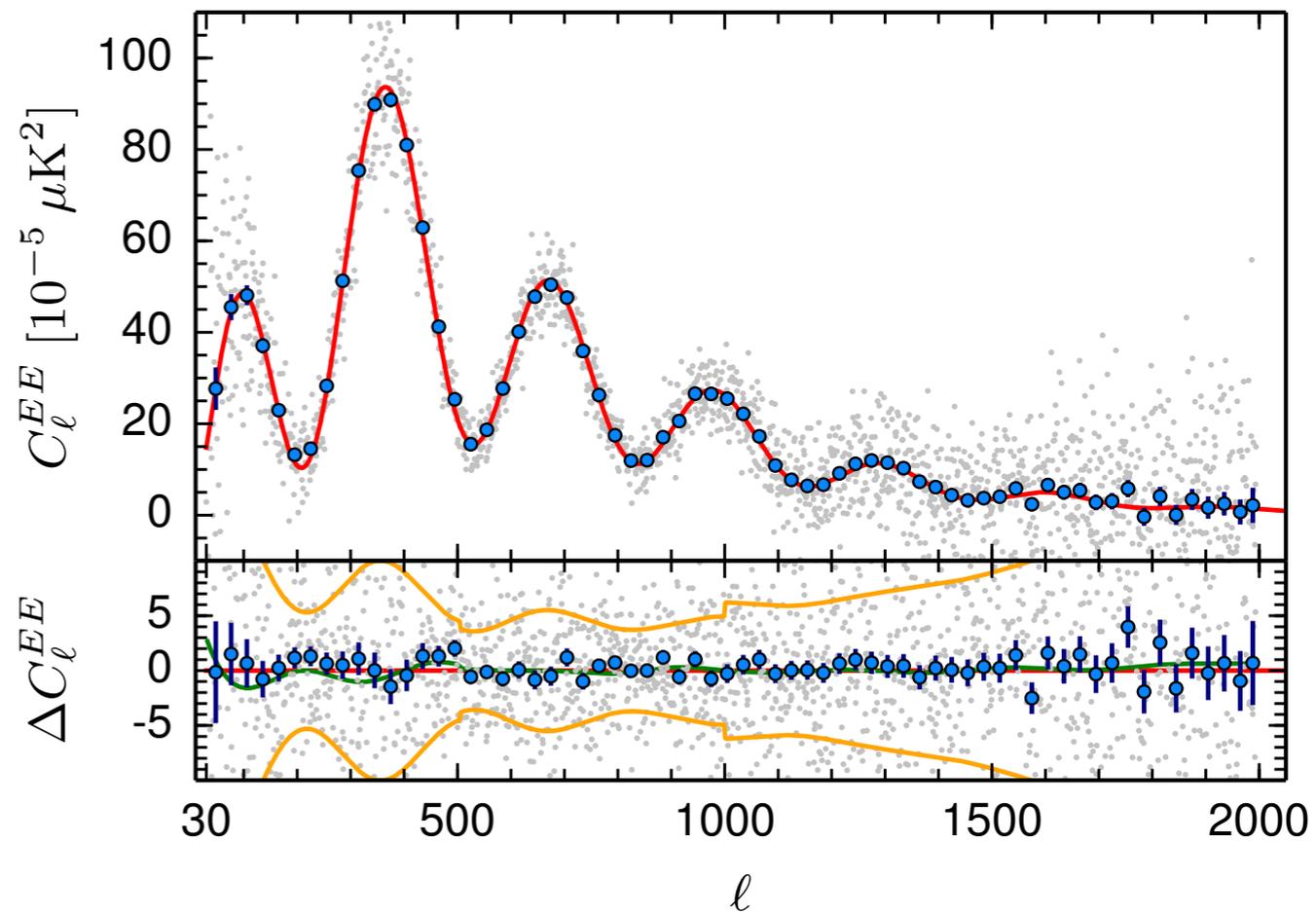


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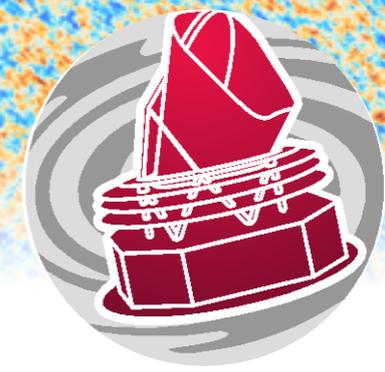
TE



EE



[Planck Collaboration XIII, *Planck* 2015 results]



- **6 minimal parameters**

- 2 for the primordial matter spectrum

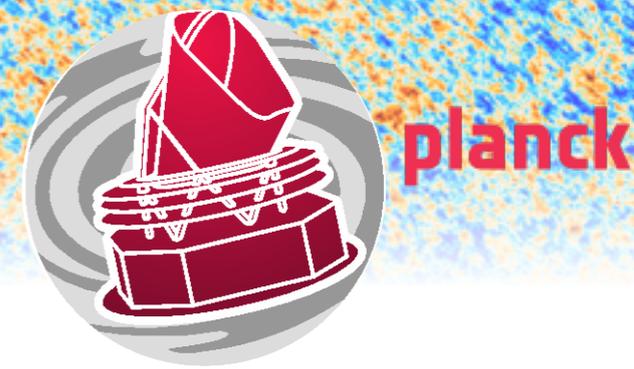
$$\mathcal{P}_{\mathcal{R}}(k) = A_s \left(\frac{k}{k_0} \right)^{n_s - 1}$$

- 1 expansion rate H_0

- 2 parameters for densities Ω_b Ω_c

- reionization optical depth τ

Λ CDM results



[Planck Collaboration XIII, *Planck* 2015 results]

	WMAP	Planck 2013	Planck 2015
Ω_b	0.0462 ± 0.0010	0.0487 ± 0.0006	0.04916 ± 0.0003
Ω_c	0.2322 ± 0.0092	0.2647 ± 0.0059	0.2647 ± 0.0033
H_0	70.0 ± 2.2	67.3 ± 1.2	67.27 ± 0.66
n_s	0.972 ± 0.013	0.960 ± 0.007	0.964 ± 0.005
$10^9 A_s$	2.189 ± 0.090	2.196 ± 0.060	2.207 ± 0.074
τ	0.089 ± 0.014	0.089 ± 0.014	0.078 ± 0.019
Ω_Λ	0.721 ± 0.025	0.685 ± 0.018	0.684 ± 0.009
Ω_m	0.279 ± 0.023	0.315 ± 0.018	0.316 ± 0.009

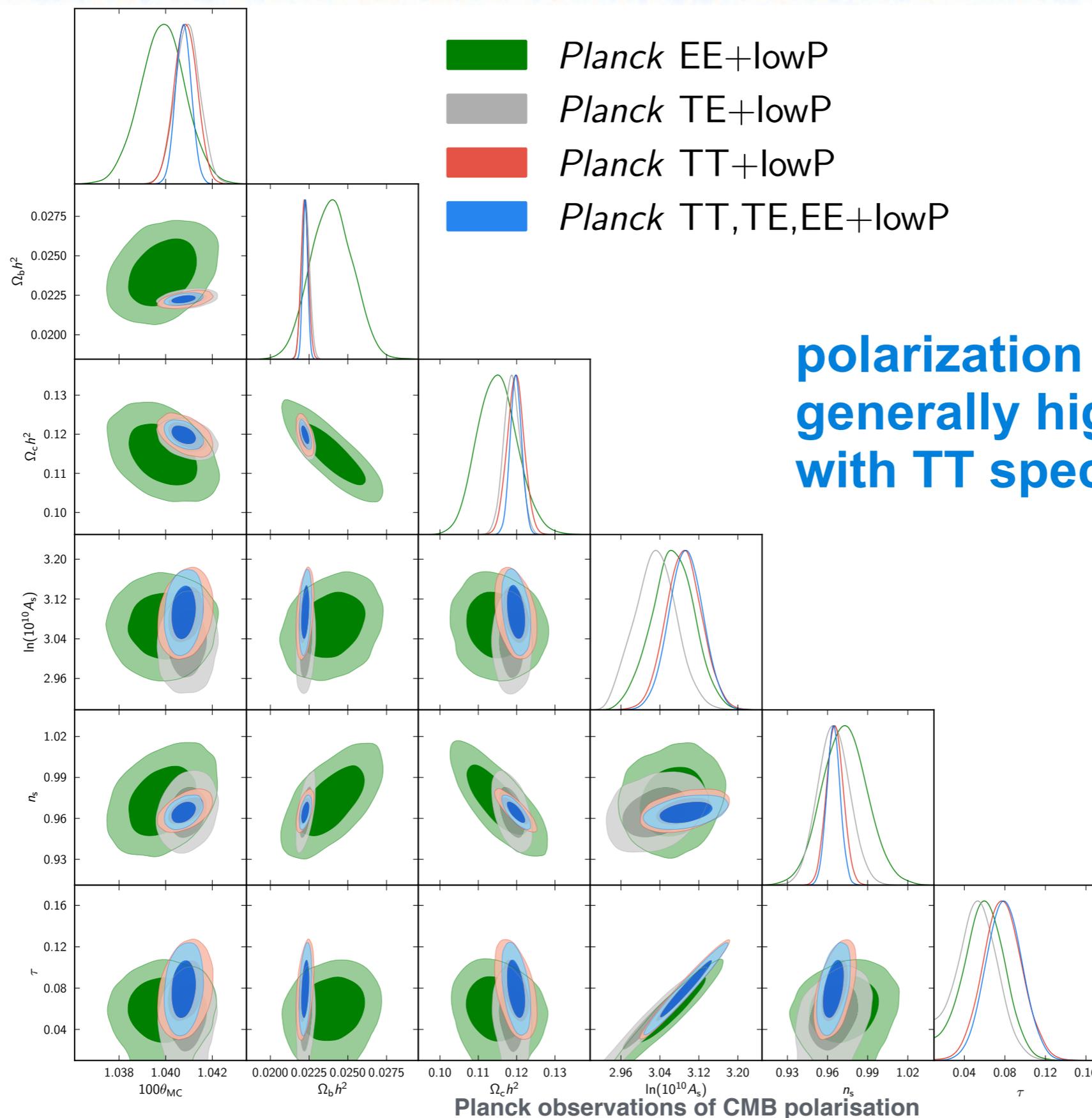
- almost factor 2 on error bars wrt WMAP
- Planck very stable between 2013 and 2015
- almost factor 2 in 2015 when adding polarization

CMB results



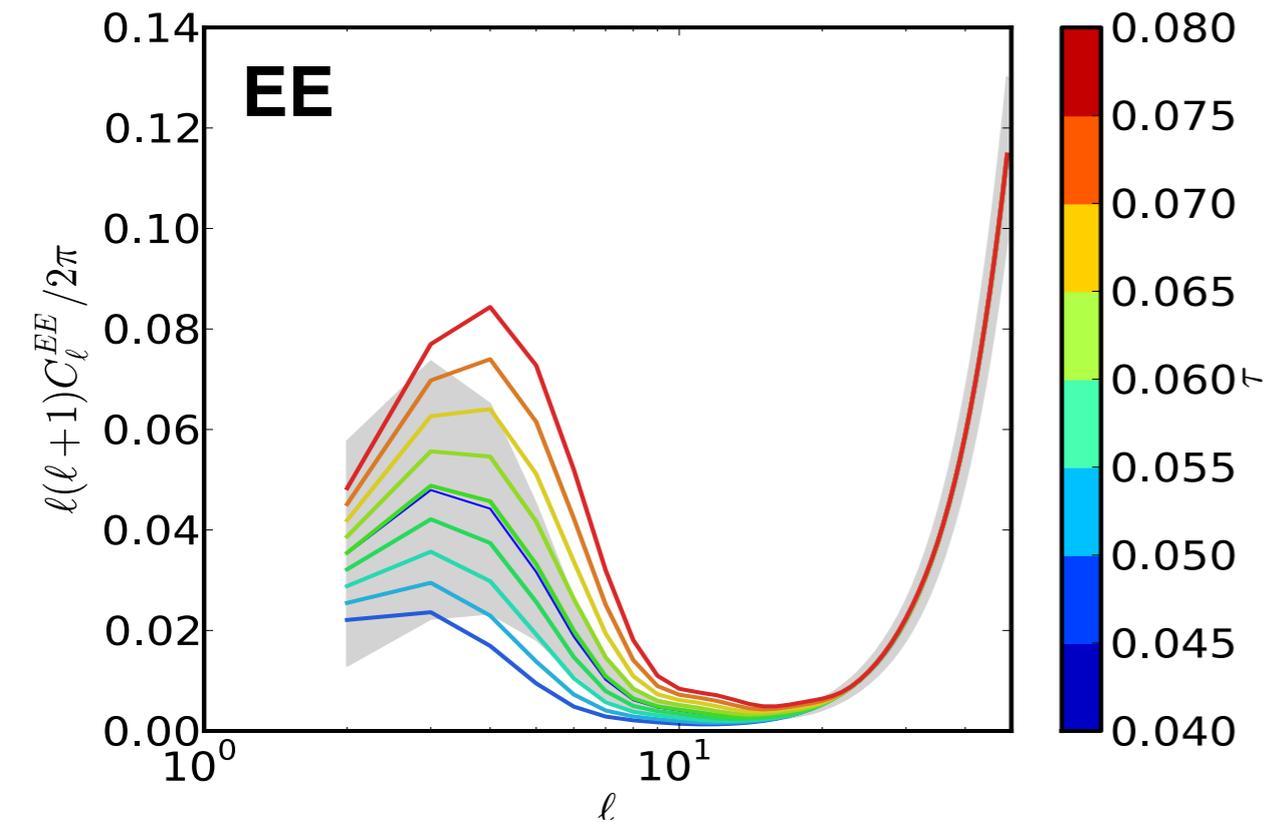
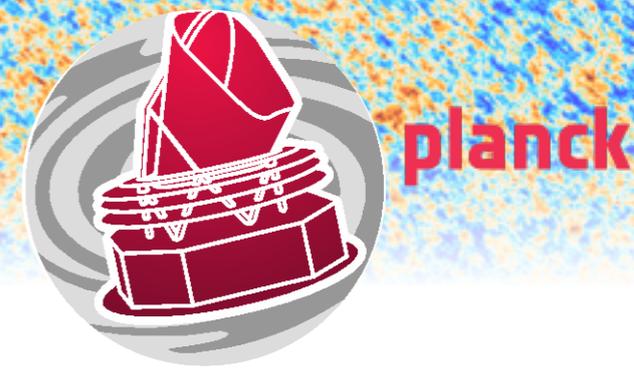
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[Planck Collaboration XIII, Planck 2015 results]



polarization spectra are generally highly consistent with TT spectra

reionisation optical depth



From CMB data:

1. WMAP 9yr

• $\tau = 0.089 \pm 0.014$

2. Planck 2013

• $\tau = 0.089 \pm 0.014$ (TT with WP)

• $\tau = 0.075 \pm 0.013$ (TT with Planck dust)

3. Planck 2015

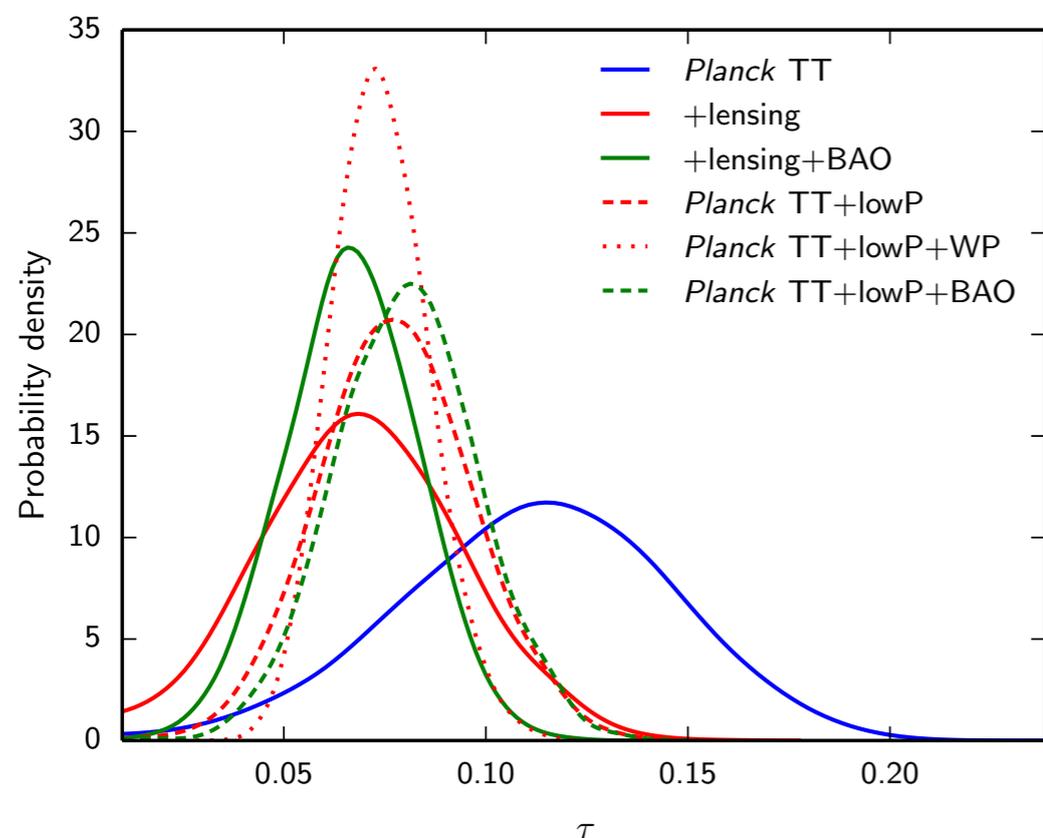
• $\tau = 0.078 \pm 0.019$ (TT + lowP)

• $\tau = 0.066 \pm 0.016$ (TT + lowP + lensing)

• $\tau = 0.067 \pm 0.016$ (TT + lensing + BAO)

4. Planck HFI polarisation EE low- ℓ

• decreasing trend continues...



[Planck Collaboration XIII, *Planck* 2015 results]

primordial gravitational waves



- **March 2014**

BICEP2 claimed 5σ detection of primordial B-modes with $r=0.2$

[PRL 112, 241101 (2014)]

- **September 2014**

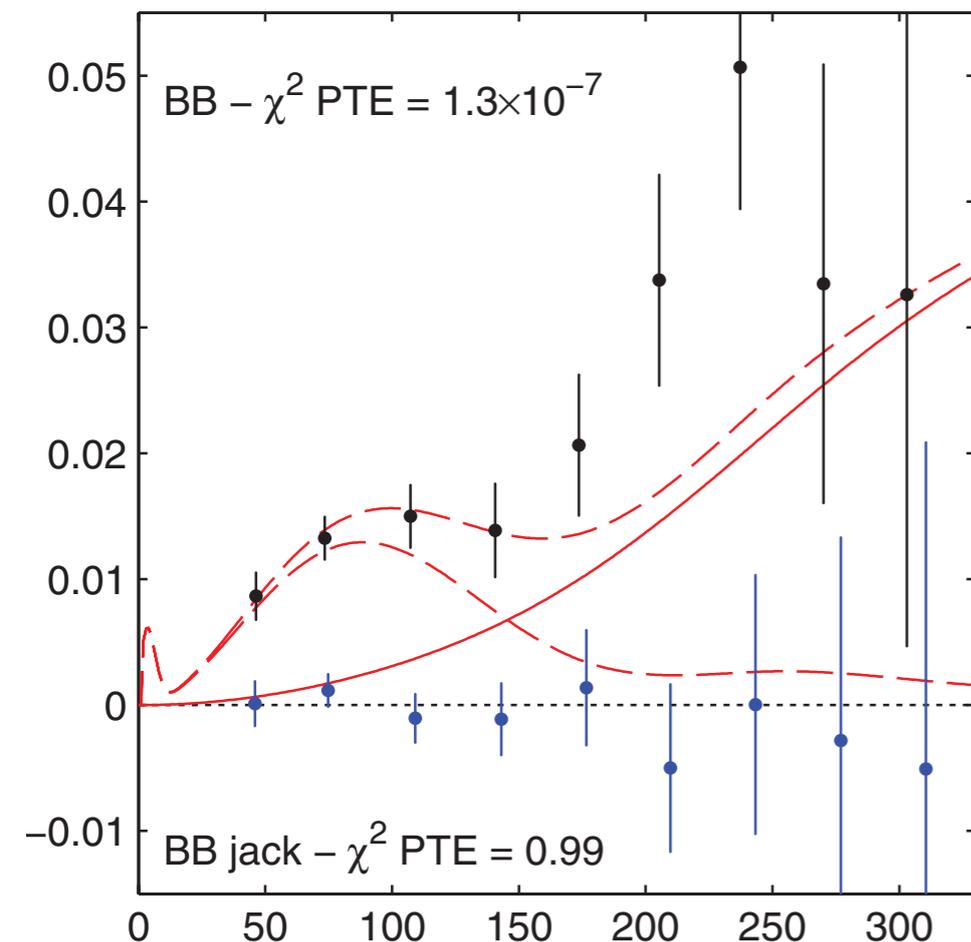
Planck showed that BICEP2 results are compatible with polarized dust emission

[Planck Intermediate XXX, arXiv:1409.5738 (2014)]

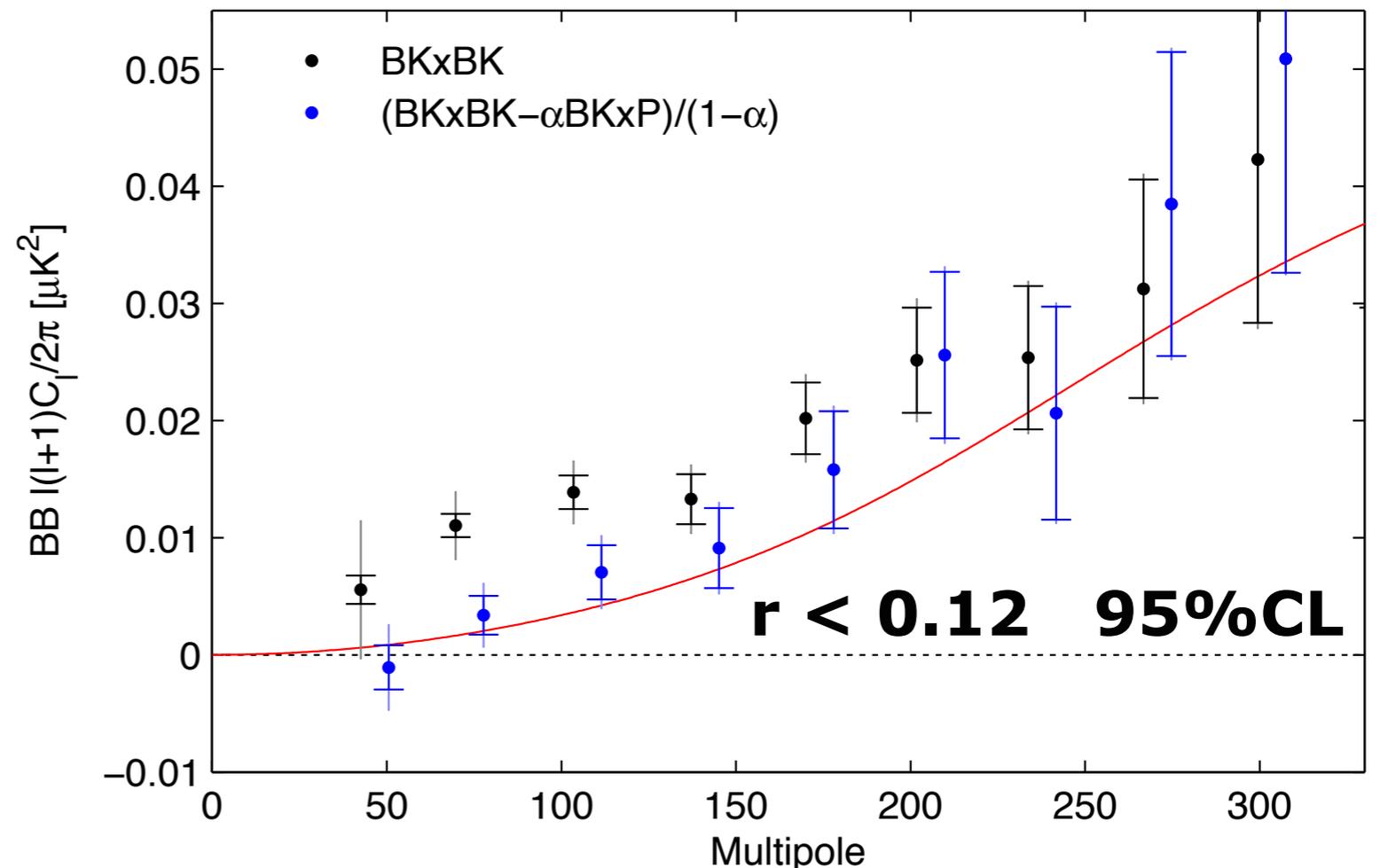
- **January 2015**

joint analysis Planck-BICEP2/Keck shows no primordial signal, give upper-limit

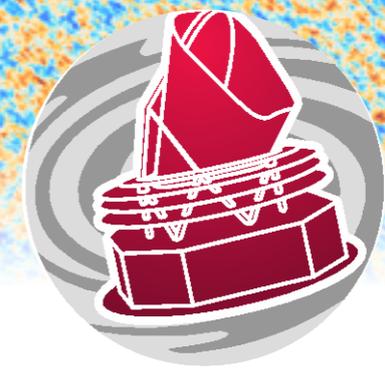
[arXiv:1502.00612 (2015)]



[BICEP2 collaboration, PRL 112, 241101 (2014)]

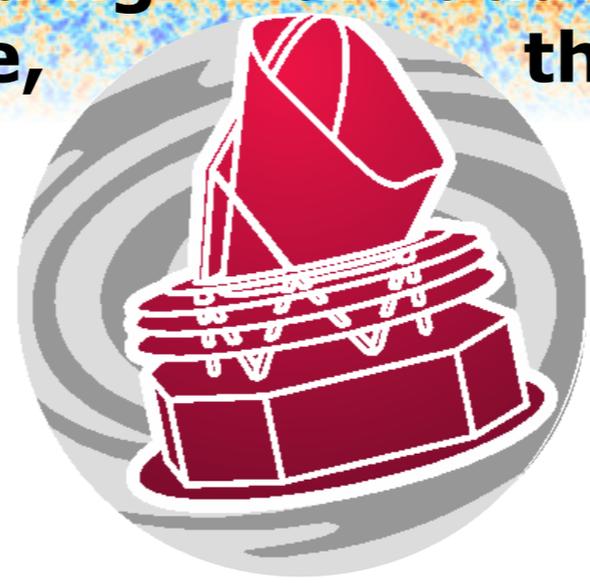


[BICEP2/Keck - Planck collaborations, arXiv:1502.00612 (2015)]



- **2015 Planck constraints now include polarization.**
- **As in 2013 Λ CDM model continues to be a good fit to the Planck data, including polarization.**
- **No convincing evidence for any simple extensions.**
- **Specific polarisation effect:**
 - Reionization optical depth seems lower than expected which disfavor large abundances of star-forming galaxies beyond $z=15$
 - Planck constraints on gravitational waves
 - $r < 0.12$ 95% CL with BICEP2/Keck (direct measurement)

The scientific results that we present today are a product of the Planck Collaboration, including individuals from more than 100 scientific institutes in Europe, the USA and Canada



Planck is a project of the European Space Agency, with instruments provided by two scientific Consortia funded by ESA member states (in particular the lead countries: France and Italy) with contributions from NASA (USA), and telescope reflectors provided in a collaboration between ESA and a scientific Consortium led and funded by Denmark.



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The End